Fundamentals of GNSS for high-precision geodesy

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GNSS Data Processing and Analysis with GAMIT/GLOBK and track Virtual (online only) hosted by EarthScope 22–26 July 2024

http://geoweb.mit.edu/gg/courses/202407_EarthScope/ Material from R. W. King, T. A. Herring, M. A. Floyd (MIT) and S. C. McClusky (now at ANU)

Outline

- GNSS Observables:
	- GNSS data and the combinations of phase and pseudo-range used
- Modeling the observations: Aspects not well modeled
	- Multipath and antenna phase center models
	- Atmospheric delay propagation
- Limits of GNSS accuracy
	- Monument types
	- Loading (more later)
	- Orbit quality

Instantaneous positioning with GNSS pseudoranges

Precise positioning using phase measurements

- High-precision positioning uses the phase observations
- Long-session static: tracking of change in phase over time carries most of the information
- The shorter the span the more important is ambiguity resolution

Observables in data processing

- Fundamental observations
	- L1 phase = f1 × range $(\lambda = 19 \text{ cm for GPS})$ L2 phase = f2 × range $(\lambda = 24 \text{ cm for GPS})$ (L5 $\lambda = 25.5 \text{ cm}$)
	- C1 or P1 pseudorange used separately to get receiver clock offset (time)
- To estimate parameters use doubly-differenced observables
	- LC = 2.546 L1 − 1.984 L2 "ionosphere-free phase combination" (L1 cycles) [LC = 2.26 L1 1.69 L5] (cycles)
	- PC = 2.546 P1 − 1.546 P2 "ionosphere-free range combination" (meters) [PC = 2.26 P1 1.26 L5] (meters)
- Double differencing (DD) cancels clock fluctuations; LC cancels almost all of ionosphere. Both DD and LC amplify noise (use Li and L2 directly and independently for baselines ≤ 1 km)
- Auxiliary combinations for data editing and ambiguity resolution: "geometry-free combination (LG)" or "extra wide-lane" (EX-WL)
	- LG = L2 $f2/f1$ L1 (used in GAMIT)
	- EX-WL = $L1 f1/f2 L2$ (used in track)
	- Removes all frequency-independent effects (geometric & atmosphere) but not multipath or ionosphere
- Melbourne-Wubbena wide-lane (MW-WL): phase/pseudorange combination that removes geometry and ionosphere; dominated by pseudorange noise
	- MW-WL = N1 N2 = (L1 L2) ($\Delta f/\Sigma f$)(P1 + P2) = (L1 L2) 0.124(P1 + P2) $[(L1-L5)-0.145(P1+P5)]$
- With GNSS processing, other frequencies are available (and changing with time).

Modeling the observations I. Conceptual/Quantitative

- Motion of the satellites
	- Earth's gravity field (flattening effect approx. 10 km; higher harmonics 100 m)
	- Attraction of Moon and Sun (100 m)
	- *Solar radiation pressure (20 m): Different for different GNSS types*
- Motion of the Earth
	- Irregular rotation of the Earth (5 m)
	- Luni-solar solid-Earth tides (30 cm)
	- *Loading due to the oceans, atmosphere, and surface water and ice (10 mm)*
- Propagation of the signal
	- Neutral atmosphere (dry 6 m; *wet 1 m*)
	- Ionosphere (10 m but LC corrects to a few mm most of the time)
	- *Variations in the phase centers of the ground and satellite antennas (10 cm)*
- * *incompletely modeled*

Modeling the observations II. Software structure

- Satellite orbit
	- IGS tabulated ephemeris (Earth-fixed SP3 file) $[\text{track}]$
	- GAMIT tabulated ephemeris (t-file): numerical integration by arc in inertial space, fit to SP3 file, may be represented by its initial conditions (ICs) and radiation-pressure parameters;
requires tabulated positions of S
- Motion of the Earth in inertial space [model or track]
	- Analytical models for precession and nutation (tabulated); IERS observed values for pole position (wobble), and axial rotation (UT1)
	- Analytical model of solid-Earth tides; global grids of ocean and atmospheric tidal loading
- Propagation of the signal [model or track]
	- Zenith hydrostatic (dry) delay (ZHD) from pressure (met-file, VMF1, or GPT)
	- Zenith wet delay (ZWD) [crudely modeled and estimated in solve or track]
	- ZHD and ZWD mapped to line-of-sight with mapping functions (VMF1 grid or GMF)
	- Variations in the phase centers of the ground and satellite antennas (ANTEX file)

Parameter estimation

- Phase observations [solve or track]
	- Form double difference LC combination of L1 and L2 or L5 to cancel clocks & ionosphere
	- Apply a priori constraints
	- Estimate the coordinates, ZTD, and real-valued ambiguities
	- Form M-W WL and/or phase WL with ionospheric constraints to estimate and resolve the WL (N2 − N1) integer ambiguities $[autcln (or solve), track]$
	- Estimate and resolve the narrow-lane (NL) ambiguities [solve, track]
	- Estimate the coordinates and ZTD with WL and NL ambiguities fixed
		- Estimation can be batch least squares $[solve]$ or sequential (Kalman filter) $[track]$
- Quasi-observations from phase solution (h-file) $[qLook]$
	- Sequential (Kalman filter)
	- Epoch-by-epoch test of compatibility (χ^2 increment) but batch output

Limits of GNSS accuracy

- Signal propagation effects
	- Signal scattering (antenna phase center / multipath)
	- Atmospheric delay (mainly water vapor)
	- Ionospheric effects
	- Receiver noise
- Unmodeled motions of the station
	- Monument instability
	- Loading of the crust by atmosphere, oceans, and surface water
- Unmodeled motions of the satellites
- Reference frame

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Multipath is interference between the direct and a far-field reflected signal (geometric optics apply)

To mitigate the effects:

- Avoid reflective surfaces
- Use a ground plane antenna
- Avoid near-ground mounts
- Observe for many hours
- Remove with average from many days

Direct Signal

Reflected Signal

Reflection Career

Multipath contributions to observed phase for three different antenna heights (from Elosegui et al., 1995)

More dangerous are near-field signal interactions that change the effective antenna phase center with the elevation and azimuth of the incoming signal

Right: Examples of the antenna phase patterns determined in an anechoic chamber…BUT the actual pattern in the field is affected by the antenna mount

To avoid height and ZTD errors of centimeters, we must use at least a nominal model for the phase-center variations (PCVs) for each antenna type

Figures courtesy of UNAVCO

Atmospheric delay

The signal from each GNSS satellite is delayed by an amount dependent on the pressure and humidity and its elevation above the horizon. We invert the measurements to estimate the average delay at the zenith (green bar).

(Figure courtesy of COSMIC Program)

Zenith delay from wet and dry components of the atmosphere

- Colors are for different satellites
- Total delay is ~2.5 meters
	- Variability mostly caused by wet component
- Wet delay is ~0.2 meters
	- Obtained by subtracting the hydrostatic (dry) delay
- Hydrostatic delay is ~2.2 meters
	- Little variability between satellites or over time
	- Well calibrated by surface pressure

Multipath and water vapor effects in the observations

One-way (undifferenced) LC phase residuals projected onto the sky in 4-hr snapshots. Spatially repeatable noise is multipath; time-varying noise is water vapor. Red is satellite track. Yellow (positive) and green (negative) residuals purely for visual effect. Red bar is scale (10 mm).

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Monuments anchored to bedrock are critical for tectonic studies (not so much for atmospheric studies)

Good anchoring:

- Pin in solid rock
- Drill-braced (left) in fractured rock
- Low building with deep foundation

Not-so-good anchoring:

- Vertical rods
- Buildings with shallow foundation
- Towers or tall building (thermal effects)

Annual component of vertical loading

Atmosphere (purple) 2-5 mm

Water/snow (blue/green) 2-10 mm

Nontidal ocean (red) 2-3 mm

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GPS Satellite

Limits to model are nongravitational accelerations due to solar and Earth radiation, unbalanced thrusts, and outgassing; and non-spherical antenna pattern

Modeling of these effects has improved, but for global analyses remain a problem

Parametric models are used in GAMIT and these evolve with time (and version)

Quality of IGS Final Orbits 2019/08 -2020/08 20 mm = 1 ppb (https://igs.org/acc/gps only/#final)

Analysis centers now < 10 mm RMS difference. Brown line is MIT GAMIT showing impact of SRP model improvements in version 10.71

Quality of IGS Final Orbits 2023/06 -2024/06 20 mm = 1 ppb (https://igs.org/acc/gps only/#final)

Analysis centers still < 10 mm RMS difference. Brown line is MIT GAMIT, consistently continuing to be one of the best fitting solutions. GAMIT solution is now combined GPS and Galileo solution

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Reference frames

- Basic Issue: How well can you relate your position estimates over time to:
	- 1. A set of stations whose motion is well modeled?
	- 2. A block of crust that allows you to interpret the motions?
- Implementation: How to use the available data and the features of GLOBK to realize the frame(s)
- Both questions to be addressed in detail in later lectures

Effect of Orbital and Geocentric Position Error (Uncertainty)

- High-precision GNSS is essentially relative!
- Baseline error (uncertainty) ~ Baseline length × geocentric SV or position error SV altitude
- SV errors reduced by averaging:
	- Baseline errors are $\sim 0.2 \times$ orbital error/20,000 km
	- e.g. 20 mm orbital error = 1 ppb or 1 mm on 1000 km baseline
- Network ("absolute") position errors less important for small networks
	- e.g. 5 mm position error \sim 1 ppb or 1 mm on 1000 km baseline
	- 10 cm position error \sim 20 ppb or 1 mm on 50 km baseline
- But SV and position errors are magnified for short sessions

Summary

- GNSS observables
	- GNSS data and the combinations of phase and pseudo-range used
- Modeling the observations: Aspects not well modeled are
	- Multipath and antenna phase center models
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- Limits of GNSS accuracy
	- Monument types
	- Loading (more later)
	- Orbit quality: Since 2000 less than 40 mm, corresponding to 2 ppb
		- Hard to improve on the IGS orbits